# Space Interferometry Mission Overview

Renaud Goullioud

MAM Testbed Manager
Flight Combiner System Engineer

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#### Outline



- SIM project
- Planet Detection
- Astrometry
- Architecture
- Status
- Technology plan
- Major testbeds

Space Interferometry Mission







# Project Background



- The foundation of SIM science objectives is tied to the 1991 National Research Council astrophysics decadal report chaired by Bahcall.
  - "Astrometry which is concerned with the measurements of the celestial sources ranks among the oldest and most fundamental branches of astronomy and now is on the verge of a technological revolution".
  - The Bahcall report recommended an Astrometric Interferometry Mission (AIM) as a high priority mission for the '90s with the following attributes:
    - "Measure positions of widely separated objects to a visual magnitude of 20 with precision of 30 micro arcseconds [µas];...a more challenging goal would be 3 micro arcseconds.
    - "...search for planets around stars as far away as 500-light years (150pc)
    - "...trigonometric determination of distances throughout the galaxy
    - "AIM would demonstrate the technology required for future interferometry missions."
- SIM re-confirmed in the 2000 NRC decadal report co-chaired by McKee and Taylor:
  - Strongly suggested a 10 µas maximum for Wide angle astrometry (rather than the 30 uas in the Bachall report).







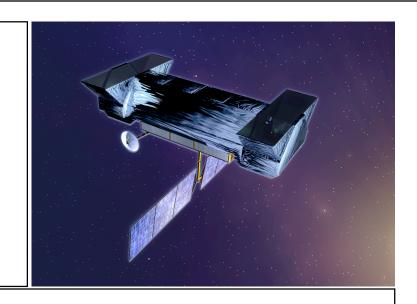






#### **Salient Features**

- 3 parallel Michelson Stellar Interferometers
- 10 meter baseline
- Visible wavelength
- Launch Vehicle: Atlas V or Delta IV ELV
- Earth-trailing solar orbit
- 5 year mission life with 10 year goal
- SIM is a JPL, Caltech, Lockheed Martin, KSC, NGST, and SIM Science Team partnership



#### **Science Goals**

- Perform a search for other planetary systems by surveying 2000 nearby stars for astrometric signatures of planetary companions.
- Survey a sample of 200 nearby stars for orbiting planets down to terrestrial-type masses.
- Improve best current catalog of star positions by >100x and extend to fainter stars to allow extension of stellar knowledge to include our entire galaxy.
- Study dynamics and evolution of stars and star clusters in our galaxy to understand how our galaxy was formed and how it will evolve.
- Calibrate luminosities of important stars and cosmological distance indicators to improve our understanding of stellar processes and to measure precise distance in the distant universe.

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#### Planet Detection



- SIM astrometry complements other indirect methods of planet detection:
  - Determines the parameters of the orbit of a planet.
  - Determines mass, the most fundamental parameter of a planet.
- Focus on  $\sim 250$  stars like the Sun (F, G, K) within 10 pc.
  - Sensitivity limit of  $\sim$ 3 M<sub>Earth</sub> at 10 pc requires 1 µas accuracy.
  - Is more sensitive than Radial Velocity (3 vs 30 M<sub>Earth</sub>) with no sin(inclination) ambiguity.
- SIM will determine the architecture of solar systems, telling us whether our solar system is rare or common.
  - Are planetary systems like our own common?
  - What is the distribution of planetary masses?
  - Are there low-mass planets in 'habitable zone'?
- SIM will also sample 2000 stars within  $\sim$ 25 pc at 4 µas accuracy.
- SIM targets stars within 25 pc that are suitable for follow-up by TPF.











# SIM Complements and Paves the Way for TPF

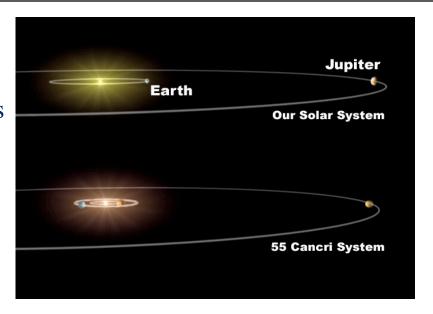


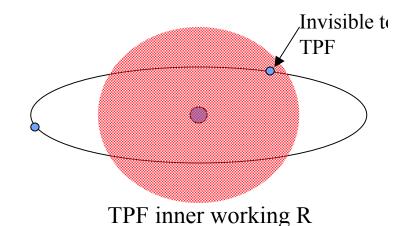
SIM will tell TPF what stars are likely to be hospitable to terrestrial planets.

Presence of Jovian planets in the wrong orbits will preclude stable orbits in the habitable zone.

- SIM's orbital information will determine when planets in eccentric/inclined orbits will be at an elongation suitable for direct detection.
- Combination of SIM masses with TPF spectroscopy of hundreds of planets will lead to new era in comparative planetology.

For stars where SIM doesn't detect a planet, and subsequently, TPF does detect a planet, SIM archival data can determine or constrain the mass of that planet with  $\sim 0.5$  Earth mass accuracy.





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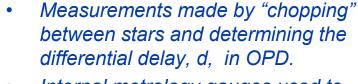
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SIM Overview



### Measurements with SIM (Interferometer Sensor)

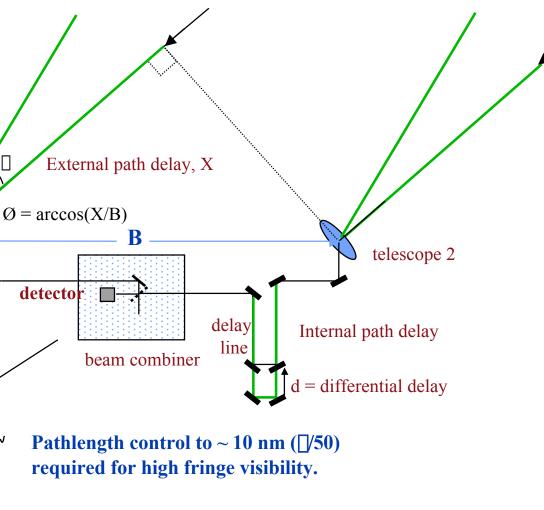




Internal metrology gauges used to measure internal path delay.

 Many such measurements are solved simultaneously to obtain stellar positions and other instrument parameters.

telescope



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f = fringe position on detector

- The peak of the interference pattern occurs at zero OPD to star









detected

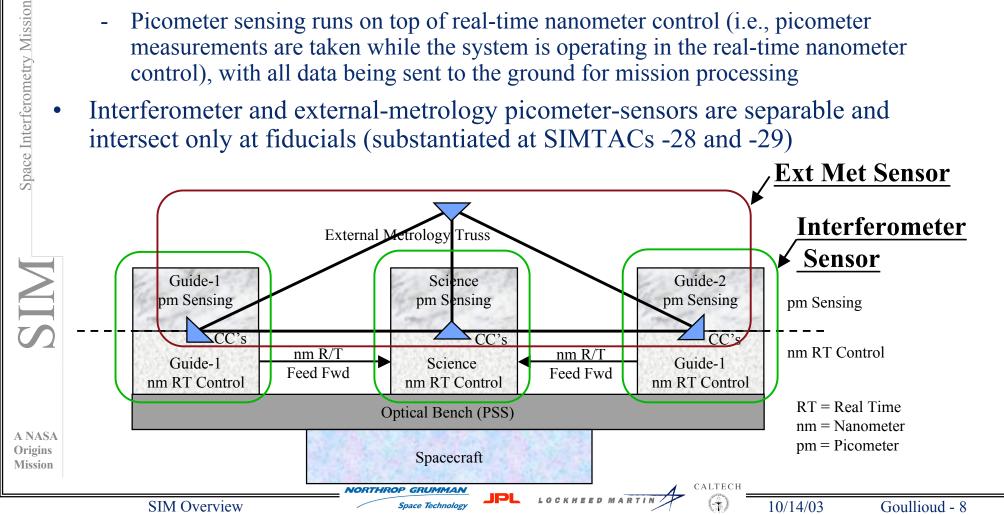
intensity



### SIM Building Blocks



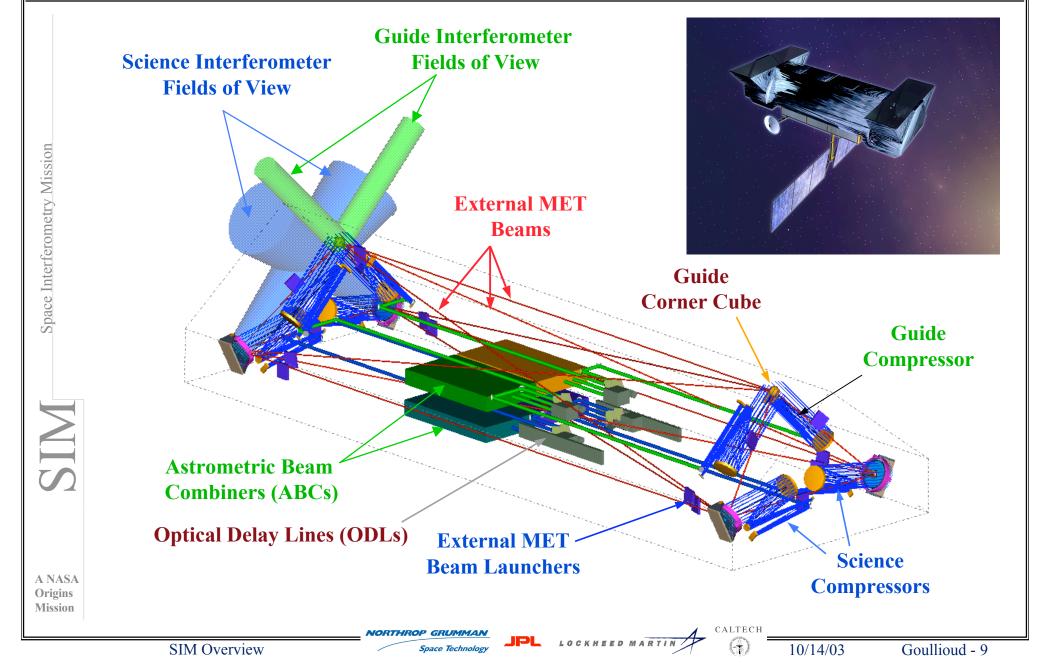
- SIM operates in two distinct regimes: (1) real-time nanometer control and (2) picometer sensing
  - Real-time nanometer control is system-wide and does not depend upon picometer sensing (nm-level external metrology information is used in real-time control)
  - Picometer sensing runs on top of real-time nanometer control (i.e., picometer measurements are taken while the system is operating in the real-time nanometer control), with all data being sent to the ground for mission processing
- Interferometer and external-metrology picometer-sensors are separable and intersect only at fiducials (substantiated at SIMTACs -28 and -29)





### SIM Instrument Configuration

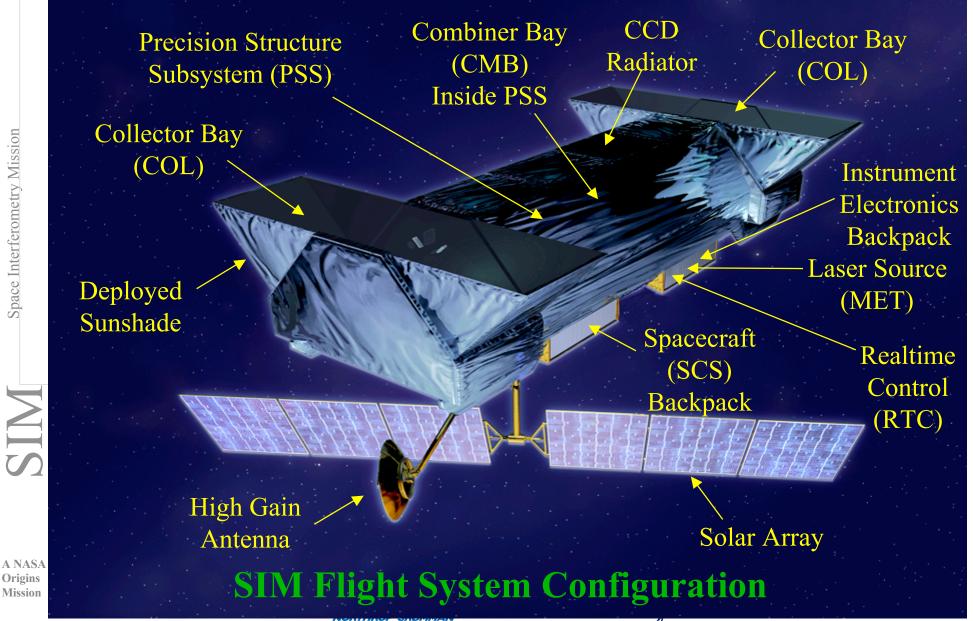






# Flight System Architecture (1)





SIM Overview



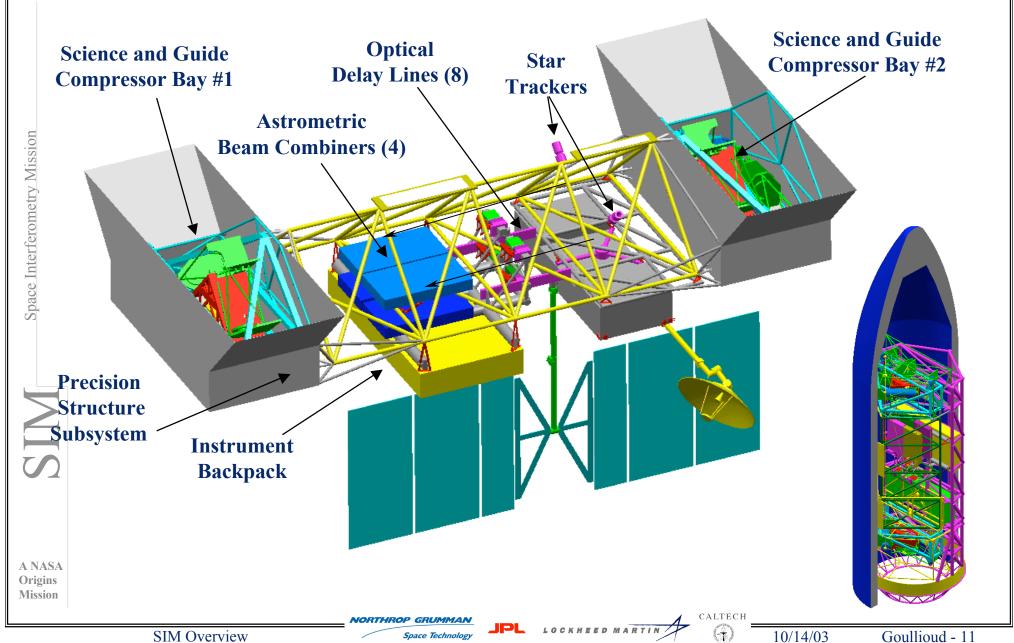






# Flight System Architecture (2)





#### Mission Status



- NASA Project Development Phases
  - Pre-Phase A: Advanced Concept Studies
  - Formulation:
    - Phase A: Mission Requirements Definition
    - Phase B: Preliminary Design
  - Implementation:
    - Phase C: Design & Development
    - Phase D: Integration & Test
      - Begins with "...the start of I&T of the full up system..."
    - Phase E: Mission Operations
- SIM has just recently passed through the NASA development Phase A to B gate
  - A very significant milestone that we've been working towards for over two decades (with just under six of the last years of that time being in Phase A).
  - We are actively preparing for the System Requirement Review in November.











### SIM Synoptic Schedule

#### August 18th Phase B start



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FY'98	FY'99	FY'00	FY'01	FY'02	FY'03	FY'04	FY'05	FY'06	FY'07	FY'08	FY'09	FY'10
Select Part			3/01 ERB		6/10/03 103 ICRR ISR	6/18/03 ICR						
	Phase A	IA Science AO Sele	no (Complete)	_	complete)	8/03	6/05 CRR 5/05 PDR/ NAR V	6/05 / CR (NASA PMC)		1/08	8/09 Ship to	12/09
			New Beam Launcher Demo (Complete)	Nano technology (STB-3) (Complete)  MAM and Kite at Level 1 (Complete)	MAM WA Basic Req (Complete) MAM NA Goal Benchmark (Com	MAM WA Goal Benchmark (2/04)	ation (4/05)	9/0 CD 10/05	R 7	se C/D	KSC	Launch
	Tech. Dev	y.	New B	Nano t	MAM	MAM Bench	Pathler Verific			3-year P tarts in (		
		e-Phase ch Gate		e-Phase ech Gate						nds in 0		

ATLO = Assy, Test & Launch Ops CA = Confirmation Assessment

CR = Confirmation Review (NASA PMC)

ERB = External Review Board PMSR = Preliminary Mission & Systems Review SRR = System Requirements Review IA = Independent Assessment

NAR = Non Advocate Review

ICRR = Initial Confirmation Readiness Review (JPL GPMC)

CDR = Critical Design Review

I&T = Integration & Test

PDR = Preliminary Design Review

ICR = Initial Confirmation Review (Code S EPMC)

NASA PMC = Programmatic Management Council

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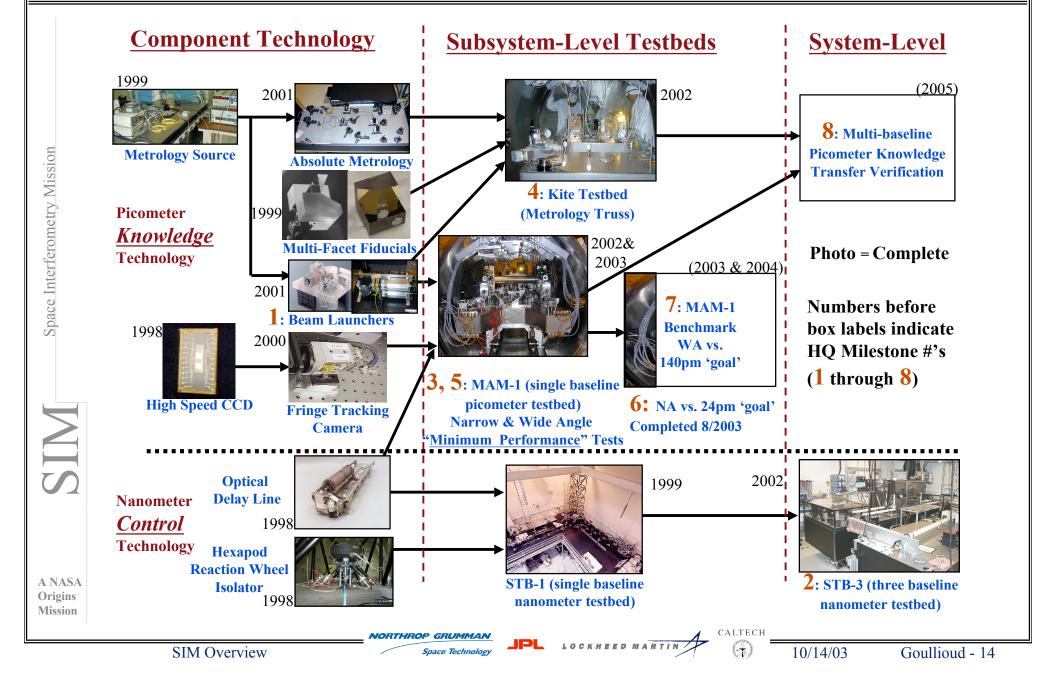






# Technology Development Diagram







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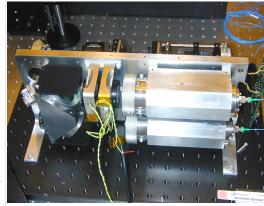
# MAM testbed (Science/Guide interferometer)

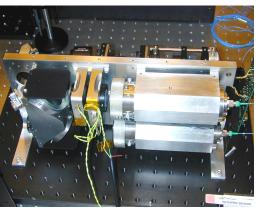


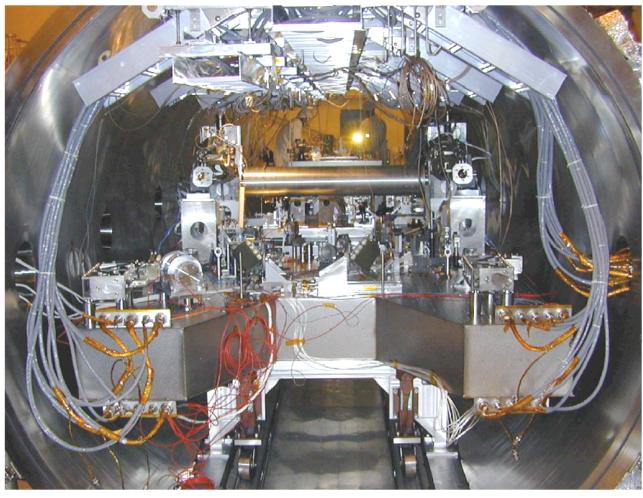


MAM Vacuum Chamber

#### **Internal Metrology** Beam Launcher







- 24 picometers agreement between internal metrology and white light fringes in the Narrow Angle field of view.
- 280 picometers agreement in the Wide Angle field of view.











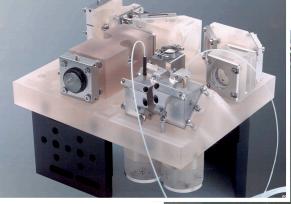
# Kite testbed (external metrology truss)

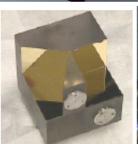


#### **External Metrology** Beam Launcher

8 picometers per metrology gauge agreement within the truss in the Narrow Angle field of view.

140 picometers agreement for Wide Angle.





Triple Corner Cube

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# System Testbed 3 (Dynamics and Control)



- Pathlength Feed-Forward demonstration:
- Pathlength stabilization of the science interferometer fringes to 30 nanometers rms open loop, using only the guide interferometers.
- 80 dB rejection of the spacecraft attitude.



3 baseline interferometer

9-meter flight like structure



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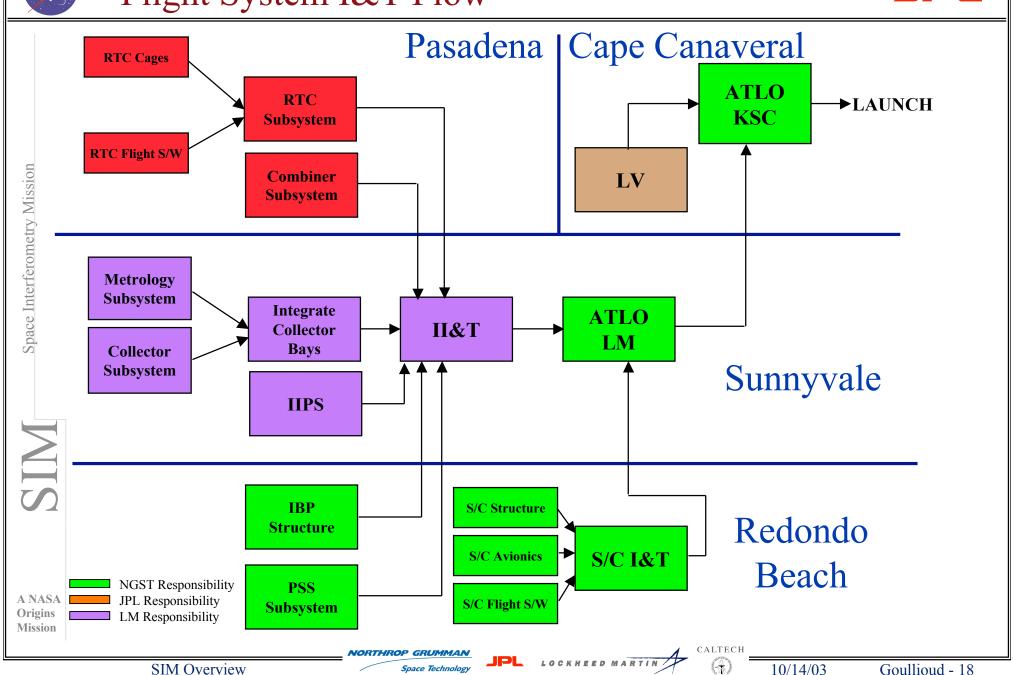






### Flight System I&T Flow





#### Conclusion



- SIM will observe about 250 stars like the Sun within 10 pc to detect the presence of planets down to 3 Earth masses at 10 pc.
- SIM will tell TPF what stars are likely to be hospitable to terrestrial planets.
- SIM's orbital information will determine when planets in eccentric/inclined orbits will be at an elongation suitable for direct detection.
- SIM technology development is quite mature.
- SIM has a stable design for the flight system.
- SIM just entered phase B.
- Scheduled launch is December 2009.

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